

# Rapidity dependence of deuteron production in central Au+Au collisions at $\sqrt{s_{NN}} = 200$ GeV

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We have measured the distributions of protons and deuterons produced in the 20% most central Au+Au collisions at RHIC ( $\sqrt{s_{NN}} = 200$  GeV) over a very wide range of transverse and longitudinal momentum. Near mid-rapidity we have also measured the distribution of anti-protons and anti-deuterons. We present our results in the context of coalescence models. In particular we extract the “volume of homogeneity” and the average phase-space density for protons and anti-protons. Near central rapidity the coalescence parameter  $B_2(p_T)$  and the space-averaged phase-space density  $\langle f \rangle(p_T)$  are very similar for both protons and anti-protons. For protons we see little variation of either  $B_2(p_T)$  or the space-averaged phase-space density as the rapidity increases from 0 to 3. However, these quantities depend strongly on  $p_T$  at all rapidities. These results are in contrast to data from lower energy collisions where the proton and anti-proton phase-space densities are different at  $y=0$  and both  $B_2$  and  $\langle f \rangle$  depend strongly on rapidity.

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## I. INTRODUCTION

Deuterons detected in heavy ion collisions are conventionally thought to be produced predominantly via a process called coalescence. Protons and neutrons that are close enough in phase-space, i.e. in position and momentum space “coalesce” to form deuterons. The surround-

ing medium created in the A+A collisions enables this ‘2 → 1’ process to proceed while conserving energy and momentum. There is a wide range of evidence to suggest that a dense system of strongly interacting partons is created in heavy ion collisions at RHIC energies [1]. In the hot and dense system produced in high-energy ion collisions, the coalescence of nucleons into deuterons cannot even begin before the the partons have frozen out into hadrons. Even then, the low binding energy of deuterons ( $\sim 2.2$  MeV), ensures that only deuterons produced late will survive to reach the detectors. Thus our sample is dominated by deuterons produced close to thermal freeze out, where the nuclear density is low [2, 3].

Early coalescence models assumed that the phase-space density of clusters is proportional to the product of the phase-space densities of individual nucleons that coalesced into them and that the momentum of the deuteron is trivially the sum of the nucleon momenta [4–6]. More recent coalescence models add insight into the nature of the phenomenon, but still relate cluster production to the product of phase-space densities [7]. For deuterons

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this is written as:

$$E_d \cdot \frac{d^3 N_d}{dp_d^3} = B_2(\mathbf{p})(E_p \cdot \frac{d^3 N_p}{dp_p^3})^2, \quad (1)$$

where  $\mathbf{p}$  is the momentum of the proton and  $B_2$  is defined as the coefficient linking the square of the nucleon distribution to that of the proton distribution.

Ideally for a coalescence study one would measure the spectra of both protons and neutrons. At  $\sqrt{s_{NN}} = 4.9$  GeV the  $n/p$  ratio measured in the 10% most central Au on Pb collisions has a value of  $1.19 \pm .08$  and is independent of  $m_T$  in a wide range of rapidity (1.6 to 2.4) [8], whereas the same ratio for the incident nuclei is equal to 1.52. (Measurements by other experiments imply that most of the isospin lost by the neutrons goes into an excess of  $\pi^-$  over  $\pi^+$  [9, 10]) At higher energies we expect  $n/p$  to be close to 1 over a wide rapidity range.

In thermodynamic models that assume thermalized distributions of nucleons,  $B_2$  carries information about the effective volume (in coordinate space) of the nucleons when they coalesce. This is also known as the homogeneity volume, and is defined as that volume over which nucleons, are close enough in momentum that they can coalesce [4–6]. At RHIC the volume of homogeneity of the emitting source has been studied in great detail at central rapidity but few measurements are available in the forward region. Consequently the deuteron coalescence analysis gives BRAHMS a unique opportunity to study the volume of the proton source over a wide range of rapidity.

The paper is organized as follows. This introduction continues with a short review of the homogeneity volume and the phase-space density averaged over coordinate space. Section II details the detector setup, the selection of the central events used for this analysis, a brief discussion about tracking, particle identification and the feed down corrections applied to the proton and anti-proton spectra. In Section III we describe the fully corrected spectra and the inferred values of  $B_2$  and average phase-space density as functions of transverse momentum and rapidity. Finally, we summarize our results and the physics that we extract from them.

### A. $B_2$ and the Homogeneity Volume

The link between  $B_2$  and the homogeneity volume can be seen in coalescence models that use the so-called sudden approximation technique to write the phase-space density of deuterons as an overlap of the nucleon wave functions with the deuteron wave function. The volume of the nucleon system is introduced through the normalization of their wave functions [11]. One can make a reasonable approximation of the deuteron's wave function by assuming it is a Gaussian of width  $\delta = 2.8$  fm. If we further assume that the region where coalescence occurs also has a Gaussian spatial profile with width  $R_G$ ,

one can write [12]:

$$(R_G^2(\mathbf{p}) + \frac{\delta^2}{2})^{3/2} = \frac{3}{2} \cdot \frac{\pi^{3/2} \hbar^3}{B_2(\mathbf{p}) \cdot m_p c^2}, \quad (2)$$

where  $m_p$  denotes the proton mass. The left hand side of this equation just represents the convolution of the proton source size and the Gaussian representation of the deuteron wavefunction. For small systems  $\delta$  dominates over  $R_G$  and the measurements have only little sensitivity to the proton source size. One advantage of this ansatz is that it facilitates comparison to interferometry radii. We show in section III that this model works well for our data since we find consistency between our estimate of the proton source size and HBT results. At lower beam energies  $R_G$  has been found to be very consistent with radii measured from  $\pi\pi$ ,  $KK$  and  $pp$  correlations [13]. The deuteron wave-function is more accurately represented by the Hulthen form, which has an exponential tail [14]. For the  $B_2$  values reported in this paper the error in  $R_G$  from using Eq. 2 is less than 0.2 fm [13].

For a variety of ion beams ranging from He to Ar incident on nuclear targets at BEVALAC energies [15], as well as  $p+A$  fixed target experiments at FNAL [16], KEK [17] and SPS [18], the measured  $B_2$  values are independent of energy and  $p_T$  and are consistent with measurements of the spatial parameters of the deuteron wave-function. At higher A+A energies including AGS Au+Au fixed target  $\sqrt{s_{NN}}=4.9$  GeV [19–24], SPS S+S, S+Pb and Pb+Pb at  $\sqrt{s_{NN}} \approx 18$  GeV [25–27], and Au+Au ( $\sqrt{s_{NN}}=200$  GeV) at the RHIC collider [28–30],  $B_2$  decreases with energy and increases with  $p_T$ . This is consistent with the formation of deuterons in an expanding medium [31–33]. Near mid-rapidity  $B_2$  decreases by a factor of  $\approx 20$  as the center of mass energy increases from  $\sqrt{s_{NN}} \approx 1$  to 17.3 GeV [26] before flattening out at RHIC energies [34, 35]. This is similar to the behavior seen in HBT radii [36, 37].

### B. Average Phase-space Density

The average of the phase-space density  $f(\mathbf{x}, \mathbf{p})$  over the system volume at freeze-out time is a quantity, which, when compared to Bose-Einstein or Fermi-Dirac statistics, allows the assessment of the degree of chemical or kinetic equilibrium reached by the system at that stage [38, 39]. This quantity also carries information about the possible multiple-particle symmetrization effects *i.e.* pion condensates. Values greater than 1.0 would indicate the presence of quantum effects [38]. This quantity is a count of the states that the system can occupy, and as such, is used to define the contribution of particular types of particles to the overall entropy of the system. The spatial average of the phase-space density is defined as the following ratio:

$$\langle f \rangle(\mathbf{p}) \equiv \frac{\int d^3 x f^2(\mathbf{x}, \mathbf{p})}{\int d^3 x f(\mathbf{x}, \mathbf{p})}, \quad (3)$$

where the integration is carried over spatial coordinates bound by the volume of homogeneity of the system at freeze-out. The formal definition of phase-space density  $f(\mathbf{x}, \mathbf{p})$  for a particle of spin  $J$  is written as:

$$f(\mathbf{x}, \mathbf{p}) \equiv \frac{(2\pi\hbar)^3}{(2J+1)} \frac{d^6 N}{dp^3 dx^3}. \quad (4)$$

For a system in chemical equilibrium at a temperature  $T$  and chemical potential  $\mu$ ,

$$f(E) = \frac{1}{e^{(E-\mu)/T} \pm 1}, \quad (5)$$

where  $E$  is the energy and  $\pm 1$  selects bosons or fermions. For a dilute system, *i.e.*  $f \ll 1$ , Eq. 5 gives:

$$f_d \approx e^{-(E_d - \mu_p - \mu_n)/T}. \quad (6)$$

Since  $E_d = m_T \cosh(y)$ , one would expect the phase-space density to be an exponential in  $m_T$ . Note in this simple derivation, we are ignoring the collective motion of the particles. At  $\sqrt{s_{NN}} = 17.3$  GeV, it was found that strong longitudinal flow could significantly reduce the space-averaged phase-space density of pions [40]. Also at this energy the inverse slope of the phase-space density was found to increase with particle mass in a manner suggestive of transverse flow [41]. Using the fact that the deuteron has energy  $E_d = E_n + E_p$  and momentum  $\mathbf{P} = 2\mathbf{p}$ , Eq. 6 implies that

$$f_d(\mathbf{x}, \mathbf{P}) = f_p(\mathbf{x}, \mathbf{p}) \cdot f_n(\mathbf{x}, \mathbf{p}) = f_p^2(\mathbf{x}, \mathbf{p}). \quad (7)$$

To extract the average phase-space density of protons one can then replace the square term in the numerator of Eq. 3 by the phase-space density of deuterons:

$$\langle f_p \rangle(\mathbf{p}) = \frac{1}{3} (E_d \frac{d^3 N_d}{dp_d^3}) / (E_p \frac{d^3 N_p}{dp_p^3}). \quad (8)$$

Furthermore, one can make use of the assumption that deuterons are formed by coalescence and satisfy Eqs. 1 and 2 to obtain an expression for the average phase-space of protons similar to what Bertsch originally suggested for pions; dividing the spectrum by the product of the HBT radii gives [38, 39]:

$$\langle f_p \rangle(\mathbf{p}) = \frac{1}{2} \cdot E_p \frac{d^3 N_p}{dp_p^3} \frac{\pi^{3/2} \cdot \hbar^3}{R_G^3(\mathbf{p}) \cdot m_p c^2}. \quad (9)$$

The phase-space densities calculated using Eq. 8 have the expected exponential dependence in  $m_T$ . When using Eq. 9 one introduces the assumption that deuterons are produced via coalescence and that the homogeneity volume extracted from deuteron distributions is also the volume of the proton source. The calculation of the space-averaged phase-space density of protons using these assumptions is close in value to the one obtained from an

assumed thermal equilibrium and has the same behavior in  $m_T$ . This suggests that the system has evolved into chemical equilibrium at freeze-out. While chemical equilibrium has been very well established at RHIC energies for the system at mid-rapidity, these data allow us to test if the matter at forward rapidity also reaches equilibrium. To ease the comparison with lower energy data we have decided to use Eq. 8 to calculate the phase density.

The coalescence parameter  $B_2$  and the space-averaged phase-space density recast the information contained in the proton and deuteron spectra into “dynamic” and “chemical” terms. The coalescence parameter  $B_2$  can be interpreted in terms of the volume of homogeneity, which depends upon the temperature of the system and the radial flow. Indeed one confirmation that we are actually seeing coalescence would be to see if this radius was consistent with the appropriate HBT radii. This was checked at  $\sqrt{s_{NN}} = 17.3$  GeV by comparing  $R_G$  to radii extracted from  $\pi\pi$ ,  $KK$ , and  $pp$  correlation measurements [13, 42–44]. In this paper, we use the very large angular and momentum ranges of the two BRAHMS spectrometers to measure the rapidity dependence of the volume of homogeneity and space-averaged phase-space density of the (anti-)proton distributions for central Au+Au collisions.

## II. ANALYSIS

The data in this paper were collected by the BRAHMS experiment during 2004. We present proton and deuteron spectra at  $\sqrt{s_{NN}} = 200$  GeV AuAu collisions with a centrality range of 0-20%. The statistics for deuterons at high rapidity for more peripheral data is limited, so a centrality dependence analysis was not warranted. The data are analyzed in four rapidity bins:  $(-0.1, 0.1)$ ,  $(0.5, 1.2)$ ,  $(1.5, 2.5)$ , and  $(2.8, 3.2)$ . We have also measured anti-proton and anti-deuteron spectra at  $y \approx 0$  and  $y \approx 0.8$ .

### A. Detector System

The BRAHMS experimental setup consists of two movable magnetic spectrometers, the Forward Spectrometer (FS) that can be rotated from  $2.3^\circ$  to  $15^\circ$ , and the Mid-Rapidity Spectrometer (MRS) that can be rotated from  $90^\circ$  to  $30^\circ$  degrees relative to the beam line. Fast Cherenkov counters at high rapidities were used to measure luminosity, to determine the interaction vertex, and to provide a start time for time-of-flight measurements.

The MRS has two Time Projection Chambers (TPCs), TPM1 and TPM2, situated in field free regions in front and behind a dipole magnet. This assembly is followed by a highly segmented scintillator time-of-flight (TOF) wall at 4.51 m. The FS consists of 4 dipole magnets D1, D2, D3 and D4 with a total bending power of up to 9.2 Tm. The spectrometer has 5 tracking stations T1 through T5. T1 and T2 are TPCs placed in front of and after the second dipole D2. T3, T4 and T5 are drift

chambers with T3 in front of D3, T4 between D3 and D4, and T5 after D4 and just in front of the Ring Imaging Cherenkov (RICH) [45]. Details on the BRAHMS experimental setup can be found in Ref. [46].

### B. Centrality selection

The centrality of the collisions at the BRAHMS interaction region was extracted from charged particle multiplicity densities measured with a multiplicity array (MA). The MA array consists of a coaxial arrangement of Si strip detectors and scintillator tiles surrounding the intersection region [47]. The pseudo-rapidity coverage of the MA is approximately  $-2.2 < \eta < 2.2$ . The centrality selection is derived from minimum-bias trigger events, which are defined using two Zero Degree Calorimeters (ZDC), requiring energy deposit equivalent to at least one neutron in each of the two detectors and also requiring a signal in the MA in order to reject Coulomb dissociation events [46, 48]. The uncertainty in the centrality determination from our multiplicity distributions is estimated to be  $\pm 4\%$  of the size of the bin for the 0-20% bin. The fraction of the inclusive yield lost by the minimum-bias trigger is estimated to be about 4% and was corrected for.

### C. Tracking

Track reconstruction starts by finding local track segments in the TPCs and (for the forward spectrometer) in the drift chambers. These detectors are in field free regions and so these local tracks are straight lines. Local tracks on either side of a magnet are matched using the effective edge approximation generating what we call locally matched tracks. The locally matched tracks are then combined in the FS to form complete tracks. The complete tracks are refitted to extract the best measurement of the particle three momentum. Tracks in the FS are required to project through the magnet D1 onto the nominal beamline. Track quality cuts are applied for the final track selection. The momentum resolution at full field (where all data presented in this paper were taken) is:  $\delta p/p = 0.0008p$ .

The momentum distributions are corrected for tracking detector efficiency and geometrical acceptance. Tracking and matching efficiencies for each of the 5 tracking stations in the FS were calculated by constructing full tracks using just four track segments and evaluating the efficiency in the fifth station by comparing the predicted position and direction of the inter/extra-polated full track in that station with the known local segments. The local track efficiency was evaluated as function of spectrometer angle and field setting, as well as function of position and direction of the local track segments. The overall tracking efficiency is typically 60%, and is known to a few per cent. In the MRS the tracking efficiencies are de-

termined from Monte Carlo simulation using embedding techniques. The systematic error is below 4%. We correct for absorption, multiple scattering, and energy loss in the detectors using GEANT 3.21 [49]. The magnitude of these corrections on the particle yields depends on the particle momenta and the spectrometer positions, and in the FS it is about 20% with an estimated uncertainty of 2%. In the MRS the corrections are less than 8% with a small systematic error. We assume that deuteron absorption factorizes from other effects and model it as the square of the proton absorption correction (at  $p_T/2$ ) as done in [13]. More details about tracking in the MRS can be found in Ref. [50].

### D. Particle Identification

In the midrapidity region, particle identification is done using time of flight measurements, while at forward rapidities we use the RICH [45]. At  $y=0$  and 0.8 we select (anti-)protons using a  $3\sigma$  cut on  $m^2 = p^2(\frac{c^2 L^2}{t^2} - 1)$ , where  $L$  is the length of the path followed by a particle,  $t$  its time of flight and  $c$  the speed of light. For deuterons a simple  $3\sigma$  cut around the (anti-)deuteron mean mass is used. These deuteron yields are corrected for background that arises from false matching of tracks with background hits in the TOF wall. This background is determined from data in the mass region outside the deuteron peak, and contributes to the systematic error of the deuteron yields by  $\sim 5 - 10\%$  decreasing with  $p_T$ . At forward rapidities the RICH provides proton/deuteron PID separation. The ring radius ( $R$ ) in the RICH depends on the particle's velocity. For a particle of mass  $m$  in the RICH  $R \approx 9.2(1 - (p/p_{thres})^2)$  cm where  $p_{thres} = 16.1 \cdot mc$ . The PID performance is shown in Fig. 1, where the bands represent the limits used for the analysis. Because of inefficiencies about 3% of the particles moving faster than their Cherenkov threshold don't produce an identified ring. Protons and deuterons with momenta between the kaon and proton thresholds do not make a ring. The contamination from mis-identified kaons and pions are subtracted to deduce the proton spectrum in this momentum range (neglecting the small fraction of deuterons). The contamination correction to the proton spectrum is  $\sim 10\%$  at  $p_T=0.5$  GeV/ $c$  and drops to  $\sim 2\%$  at  $p_T=1.5$  GeV/ $c$ . The error on this correction is negligible. This procedure cannot be used in the region between the proton and deuteron thresholds since the contamination from pions, kaons and protons dominates the small deuteron yield.

### E. Feed-down corrections

We have corrected our data to account for the hyperons that decay into protons using the method described in

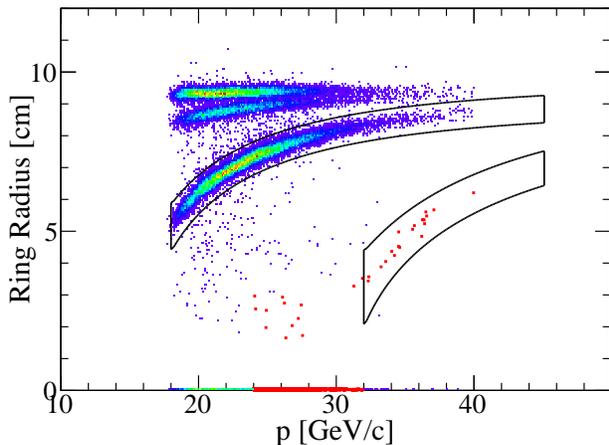


FIG. 1. (color online). Ring radius versus momentum for particles at  $y \approx 3$  showing the pion, kaon, proton and deuteron separation. The bands show the proton and deuteron selection used in the analysis.

[52]. The correction factor,  $C$ , is given by:

$$C = \frac{N_p}{N_p + N_\Lambda + N_{\Sigma^+}}, \quad (10)$$

where  $N_p$  is the number of primary protons and  $N_\Lambda$  and  $N_{\Sigma^+}$  are the number of protons coming from  $\Lambda$  and  $\Sigma^+$  decay respectively. Near central rapidity we have used the  $\Lambda$  spectra measured by PHENIX [51] and estimated the  $\Sigma/\Lambda$  ratio from lower energy measurements [52]. Since there are no measurements of  $\Lambda$ s at forward rapidities we have estimated the  $\Lambda/p$  ratio based on thermal models that were fitted to the rapidity densities of charged pions, kaons, protons and anti-protons measured by BRAHMS in the forward region [53].

With our model calculations, we find that the ratio  $\frac{dN_\Lambda}{dy} / \frac{dN_p}{dy}$  varies slowly with rapidity up to rapidities  $y \approx 4$ . The systematic error from uncertainties on the yields and the model extrapolation is estimated to be less than 3%. The correction factor also depends on the  $p_T$  dependence of the  $\Lambda/p$  ratio. BRAHMS has found that the mean transverse kinetic energy scales linearly with the mass of the particle with a slope that depends only weakly on rapidity [54]. We have used these systematics to estimate the inverse slope ( $T$ ) of the  $\Lambda$   $m_T$ -distribution. To estimate the systematic error on the  $p_T$  dependence of the correction factor we have taken the limiting cases of  $T_\Lambda = T_p$  and  $T_\Lambda = m_\Lambda/m_p T_p$ . This produces an error on the correction factor that is almost zero at  $p_T = 1$  GeV/c and reaches  $-9\%$ ,  $+6\%$  at  $p_T = 2$  GeV/c. The correction factor as a function of  $p_T$  and rapidity is shown in Fig. 2. It varies only weakly with rapidity and has a systematic error that is small in comparison to the statistical errors on  $B_2$  and the space-averaged phase-space density. At  $p_T = 2$  GeV/c the total error from the feed-down correction reaches a value of 19% for  $B_2$  and 10% for the space-averaged phase-space density.

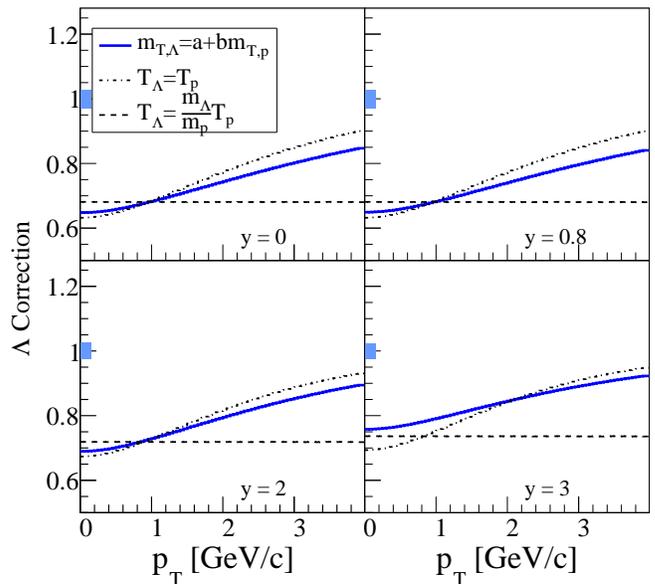


FIG. 2. (Color online). The  $p_T$  dependent feed-down correction factors for protons versus  $p_T$  and rapidity. The 3 sets of lines show the effect of various scenarios for the difference in shape between proton and  $\Lambda$  spectra. Systematic errors on the correction due to uncertainties in the yield ratio of protons and lambdas are shown by the bands to the left of the plots. The curves for anti-protons are very similar.

### III. RESULTS

Figure 3 shows the invariant proton and deuteron spectra versus  $p_T$  for the four measured rapidity bins. The systematic error on the normalization of the spectra is estimated to be about 5% with a point-to-point error in the spectra of  $\approx 8\%$  arising from the merging of spectrometer settings and PID and uncertainties in tracking efficiency determination. We fit exponentials to the spectra and extract the invariant yields,  $dN/dy$  and inverse slopes  $T$ . These are listed in Table I. About 40-50% (dependent on rapidity) of the yield of the deuterons is contained within our acceptance. For protons the fractions are slightly higher 55-60%. The proton yield falls slightly from  $y=0$  to  $y=3$ , but no firm conclusion can be made for the deuterons due to the limited  $p_T$  range at high rapidity. The inverse slopes tend to decrease with rapidity, suggesting a decrease of radial flow. The inverse slopes of the deuterons are somewhat higher than those of the protons. The average over all rapidities for the ratio of the inverse slopes is  $1.6 \pm 0.1$ . For anti-deuterons the inverse slopes are a factor of  $2.0 \pm 0.2$  higher than that of anti-protons.

Figure 4 shows  $B_2$  versus  $p_T$  and rapidity.  $B_2$  increases with  $p_T$ , which is consistent with previous experiments [26, 35]. Using Eq. 2, we find that at central rapidity the source radius  $R_G$  falls from  $4.2 \pm 0.2$  fm to  $3.1 \pm 0.4$  fm as  $m_T$  increases from 1.2 to 1.9 GeV/c<sup>2</sup>. This is consistent with the  $m_T$  dependence of HBT radii,  $R \propto 1/\sqrt{m_T}$  that

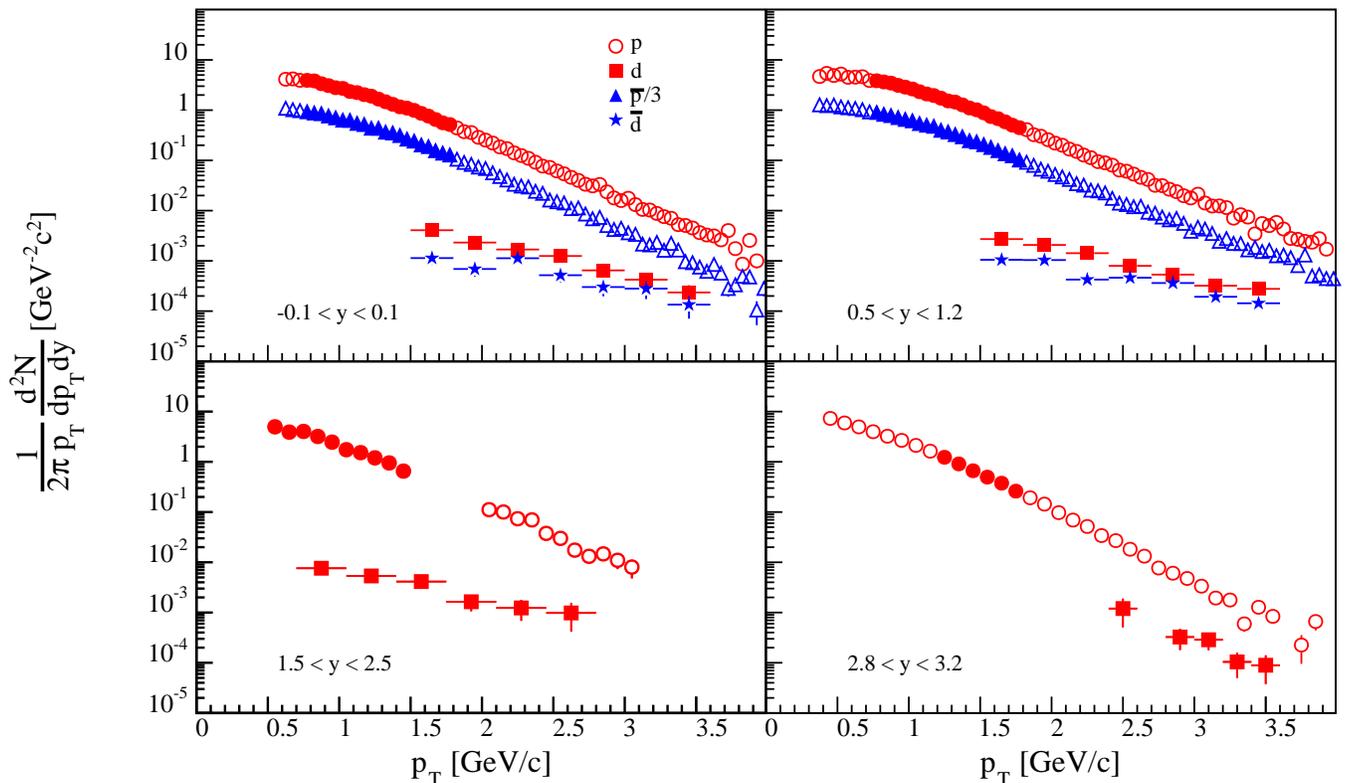


FIG. 3. (color online). (anti-)proton and (anti-)deuteron  $p_T$  spectra at various rapidities for the top 20% most central Au+Au collisions. The filled symbol part of the spectra show the  $p_T$  intervals used in the coalescence analysis. Note the limits for the deuteron range are exactly twice those used for the protons. The errors are statistical only. The horizontal bars represent the bin width.

$y$	Proton			Deuteron		
	$p_T$ fit	$dN/dy$	$T$ MeV	$p_T$ fit	$dN/dy$	$T$ MeV
0.0	0.7-4.0	$27.9 \pm 0.1$	$354 \pm 2$	1.5-3.3	$0.093 \pm 0.008$	$570 \pm 70$
		$20.8 \pm 0.1$	$352 \pm 1$		$0.033 \pm 0.004$	$870 \pm 220$
0.8	0.5-4.0	$26.0 \pm 0.1$	$356 \pm 1$	1.5-3.3	$0.068 \pm 0.003$	$610 \pm 30$
		$17.9 \pm 0.1$	$361 \pm 1$		$0.031 \pm 0.002$	$700 \pm 80$
2.0	0.5-3.0	$20.9 \pm 0.2$	$314 \pm 3$	0.8-2.8	$0.082 \pm 0.011$	$460 \pm 110$
3.0	0.5-4.0	$23.4 \pm 0.1$	$282 \pm 1$	2.5-3.5	$0.25^{+0.30}_{-0.08}$	$310 \pm 90$

TABLE I. Proton and deuteron yields,  $dN/dy$ , and inverse slopes,  $T$  (MeV), derived from fitting spectra. The lower rows at rapidities  $y = 0$  and  $y = 0.8$  are for the anti-particles. The errors listed are statistical only. For  $dN/dy$  the systematic errors due to extrapolation to low  $p_T$  is 5-10% with an additional 8% from normalization and various corrections. The systematic errors on the inverse slope are less than 10%.

has been observed by PHENIX [36] and STAR [37] and also at SPS energies [43]. The solid line in each rapidity panel represents an exponential fit to our data at  $y=0$ . We see no evidence of any rapidity dependence of  $B_2(p_T)$ . The proton and anti-proton  $B_2$  values are very close at

this energy implying a similar source size. This is in contrast to the measurements at  $\sqrt{s_{NN}} = 17.3$  GeV where the anti-proton source volume was found to be somewhat larger than the proton source [13].

Figure 5 shows the average phase-space density  $\langle f \rangle(m_T)$  for protons and anti-protons as a function of rapidity. These values are calculated using Eq. 8. The space-averaged phase-space density decreases as the  $m_T$  increases as expected from Eq. 5. The solid curve in each panel of Fig. 5 is an exponential fit to the proton density at  $y=0$ . We see little rapidity dependence of  $\langle f \rangle(p_T)$ .

From Eq. 5 we would expect the ratio of the proton and anti-proton phase densities to be flat as a function of  $p_T$ . Fitting a constant to  $\langle f_{\bar{p}} \rangle / \langle f_p \rangle$  yields a  $\chi^2/NDF$  of 5.0/6 and 13.3/6 at  $y = 0$  and  $y = 0.8$ , respectively. Near  $y=0$  the anti-protons have a slightly smaller value of space-averaged phase-space density compared to that of protons, suggesting a small chemical potential. At  $\sqrt{s_{NN}} = 17.3$  GeV the anti-proton space-averaged phase-space density was 38 times smaller [13], which suggests a much larger baryo-chemical potential at the lower energy. The inverse slope derived from the space-averaged phase-space density is  $T = 930 \pm 110$  MeV for protons. This is consistent with data at  $\sqrt{s_{NN}} = 17.3$  GeV, but much higher than the value extracted at  $\sqrt{s_{NN}} = 4.9$  GeV

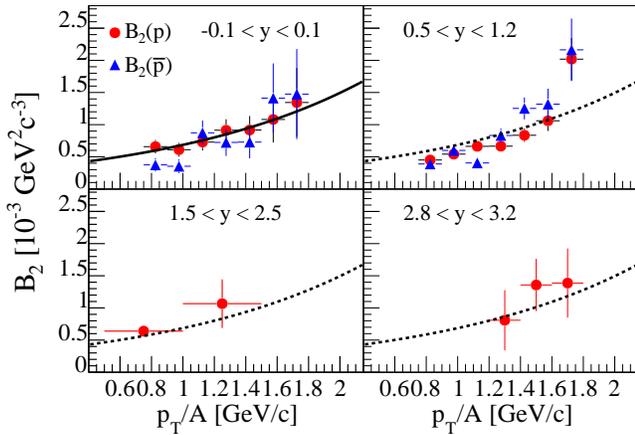


FIG. 4. (color online).  $B_2$  versus transverse momentum per nucleon at several rapidities for central Au+Au collisions at  $\sqrt{s_{NN}} = 200$  GeV. Protons are displayed with filled circles and anti-protons with filled triangles. The solid line in the top left panel is an exponential fit to the data at  $y=0$ . This same line is shown, in dotted form, in the other 3 panels. The errors are statistical only.

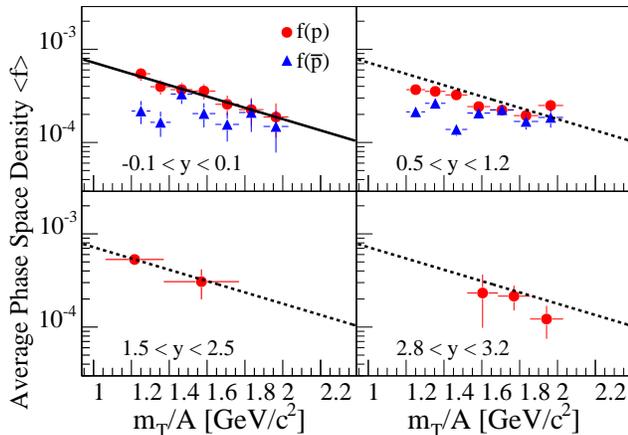


FIG. 5. (color online). The (anti-)proton space-averaged phase-space density  $\langle f \rangle(m_T/A)$  for central  $\sqrt{s_{NN}} = 200$  GeV Au+Au collisions at several rapidities. The solid line in the top left panel is an exponential fit to the data at  $y=0$ . This same line is shown, in dotted form, in the other 3 panels. The errors are statistical only.

where the inverse slope is about 350 MeV, see Fig. 6. It should be noted that the proton, kaon and pion spectra at mid-rapidity can be well described by blast-wave fits, which suggest that this increase in the inverse slope with  $\sqrt{s_{NN}}$  is largely driven by an increase in radial flow. This is supported by the fact that the phase-density of pions is characterized by a much smaller inverse slope ( $\approx 140$  MeV) than that of protons [13].

Figures 4 and 5 imply that the volume of homogeneity,  $\sim 1/B_2$ , and the space-averaged phase-space density

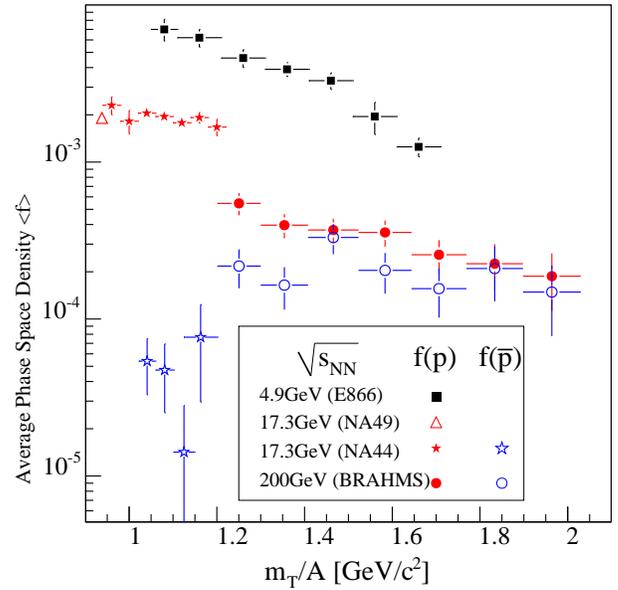


FIG. 6. (color online). The proton space-averaged phase-space density for central collisions as a function of  $\sqrt{s_{NN}}$  and  $m_T$  at  $y=0$  [13, 23, 26, 27]. The errors are statistical only.

vary little with rapidity at  $\sqrt{s_{NN}} = 200$  GeV. This is in stark contrast to the situation at lower energy. Table II shows the rapidity dependence of  $B_2$  and the space-averaged phase-space density at  $p_T = 0$  for central Pb+Pb collisions at  $\sqrt{s_{NN}} = 17.3$  GeV [13, 25, 27]. These data are compared to BRAHMS results at  $p_T/A = 1.3$  GeV/c which show very little rapidity dependence. Finally, Fig. 6 shows the evolution of the space-averaged phase-space density near mid-rapidity as the energy of the system grows from AGS to RHIC values. The space-averaged phase-space density of protons decreases with energy while that of anti-protons increases. At AGS energies  $\langle f \rangle$  at  $m_T = 0$  is 10 times bigger than its value at RHIC, while at SPS it is 2 times bigger. In contrast, the pion average phase-space density has the opposite behavior as the energy of the colliding system increases; the values at RHIC [39] are higher than those at SPS by a factor  $\sim 2$  where  $\langle f \rangle = 0.45$  [55]. At  $\sqrt{s_{NN}} = 4.9$  GeV  $\langle f \rangle$  was extracted from HBT and spectra measured at high rapidity  $\langle y \rangle \sim 3.1$  in Au+Au 10% central collisions and its estimated value at mid-rapidity is  $\sim 5$  times smaller than the RHIC value [22]. For anti-protons the growth in  $\langle f \rangle$  is similar to that of pions. The different trend for the protons may be driven mainly by baryon transport; at AGS, most of the beam protons end up at mid-rapidity, less so at SPS, and at RHIC, less than 10% of the beam protons are transported to mid-rapidity. The proton rapidity loss adds beam protons to the Gaussian rapidity density of produced protons and transforms the overall proton rapidity density into a flat distribution.

$(p_T/A = 0)$	$y=0.2$	$y=0.8$	$y=1.3$
$B_2 \times 10^4$	$7.9 \pm 0.8$	$8.1 \pm 0.3$	$13.7 \pm 2.7$
$f \times 10^3$	$1.9 \pm 0.1$	$2.5 \pm 0.2$	$3.3 \pm 0.3$

$(p_T/A = 1.3 \text{ GeV}/c)$	$y=0.0$	$y=0.8$	$y=2.0$	$y=3.0$
$B_2 \times 10^4$	$8.1 \pm 1.0$	$6.6 \pm 0.6$	$10.6 \pm 0.4$	$8.1 \pm 0.5$
$f \times 10^4$	$3.6 \pm 0.6$	$2.4 \pm 0.3$	$3.1 \pm 1.0$	$2.3 \pm 1.0$

TABLE II. The rapidity dependence of  $B_2$  and the  $f$  for (Top) Pb+Pb collisions at  $\sqrt{s_{NN}} = 17.3 \text{ GeV}$  and (Bottom)  $\sqrt{s_{NN}} = 200 \text{ GeV}$  AuAu collisions. For the 17.3 GeV data the centrality is 23% at  $y=0.2$  [27] and 20% at  $y=0.8$  and 1.3 [13, 25].

#### IV. SUMMARY

The rapidity dependence of the deuteron production in 0 – 20% central Au+Au collisions at 200 GeV has been studied in the context of coalescence models. Near central rapidity the proton and anti-proton phase-space densities are very similar, suggesting a small baryon chemical potential. The coalescence parameters, or  $B_2$ -values, for deuterons and anti-deuterons are also very close, suggesting similar freeze-out volumes for protons and anti-protons.  $B_2$  increases with  $p_T$  as expected for a system undergoing transverse flow; flow introduces a correlation between position and momentum. At a given  $p_T$  the deuteron coalescence parameter  $B_2$  is independent of rapidity, which would imply that the volume of homogeneity for protons of a given  $p_T$  is almost constant from  $y=0$  to  $y=3$ . Baryon transport in these colliding systems may be affecting the rapidity density of protons. It is interesting to note that the radial flow varies weakly with rapidity [54]. The weak dependence of radial flow on rapidity offers an explanation of why the  $p_T$  dependence

of  $B_2$  does not depend upon rapidity. It does not however address the fact that the overall magnitude of  $B_2$  is constant from  $y=0$  to  $y=3$ . The proton space-averaged phase-space density also shows no significant rapidity dependence, while depending strongly on  $m_T$  and  $\sqrt{s_{NN}}$ . The space-averaged phase-space density can be thought of as the ratio of the number of protons per unity rapidity over the volume of the proton source. The rapidity density of protons decreases by only a factor of  $0.84 \pm 0.01$  from  $y=0$  to  $y=3$ . The constancy of the space-averaged phase-space density in rapidity is consistent with the similar behavior found for  $B_2$ . Thus we observe a striking invariance of the proton source over several units of rapidity.

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